

Young Massive Clusters (YMCs)

Stellar Populations and Connection to Globular Clusters

Nate Bastian (Liverpool)

Introduction to Stellar Clusters

Pleiades



Open Clusters

few - 10^4 M_{sun}

few Myr - few Gyr

~solar metallicity

disk of the Galaxy

Globular Clusters

10^4 - 10^6 M_{sun}

10 - 13 Gyr

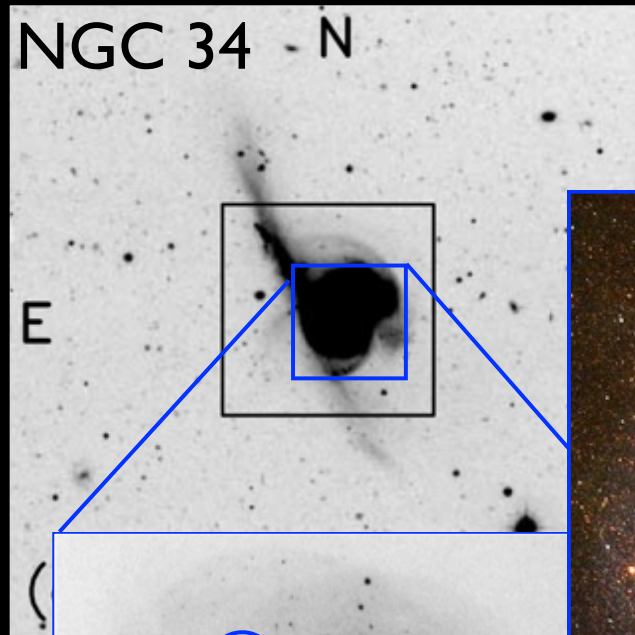
low metallicity

bulge/halo of the Galaxy

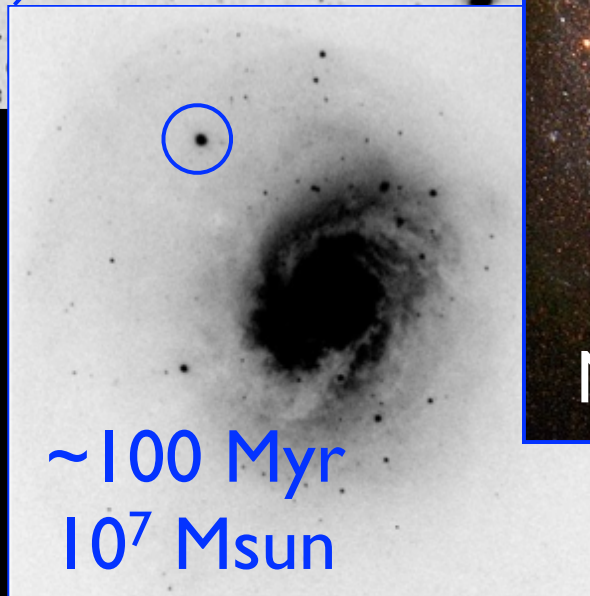
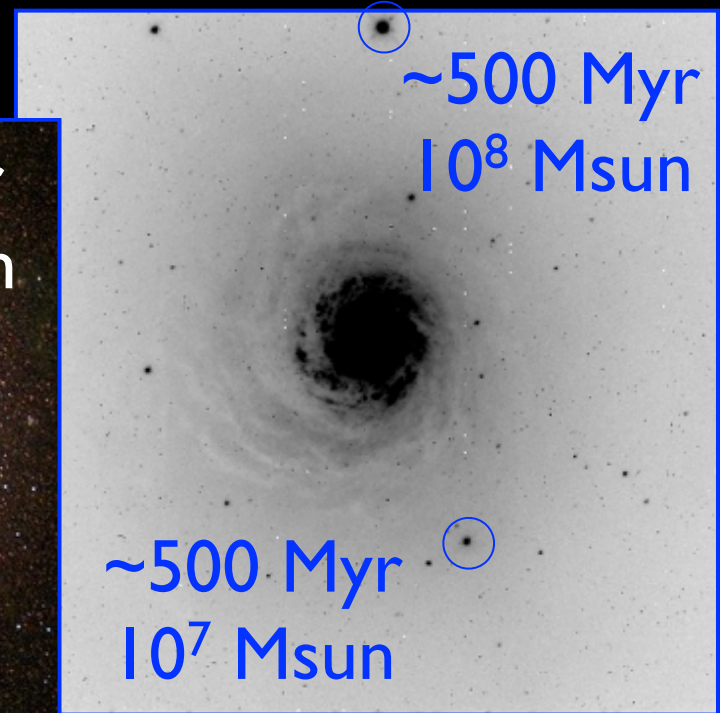
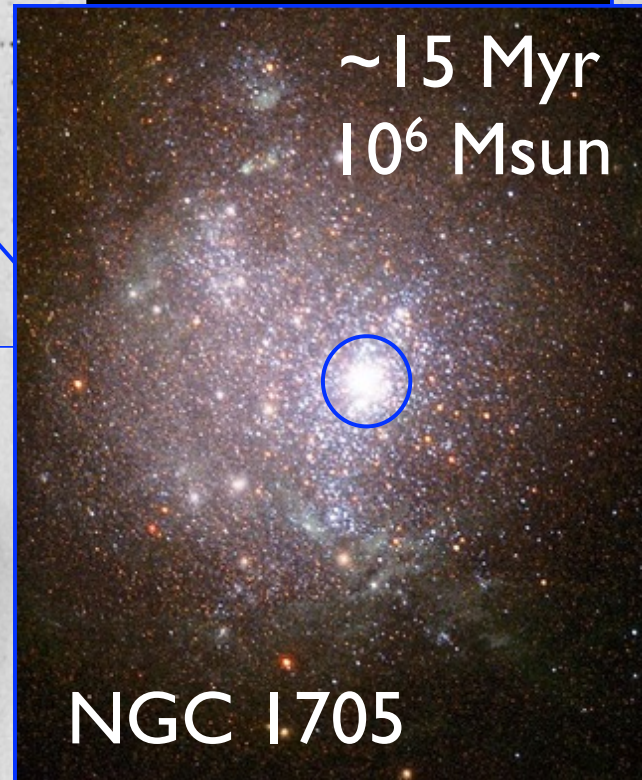


M80

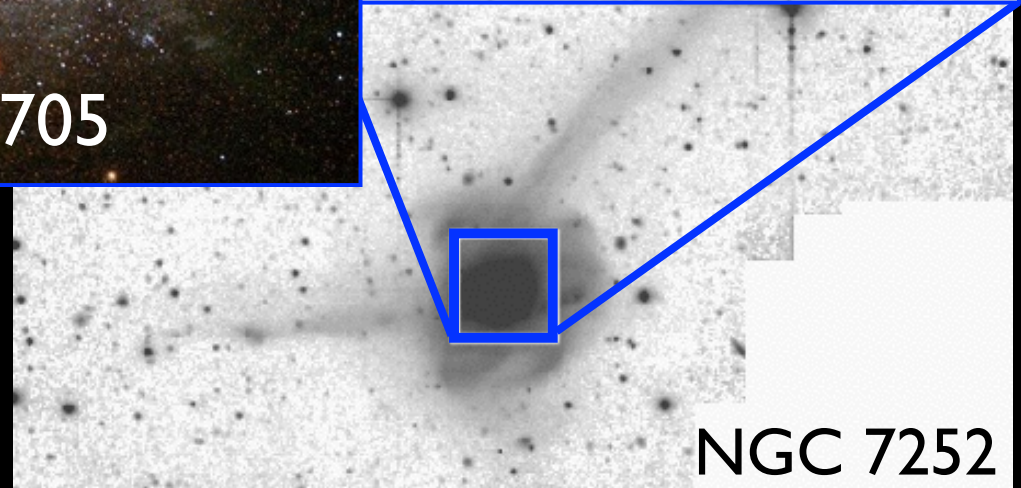
Young Massive Clusters (YMCs)



Schweizer & Seitzer 2007
Cabrera-Ziri et al. 2014



Maraston et al. 2004
Bastian et al. 2006





Omega Cen - ESO

few million M_{sun}
 ~ 12 Gyr

*globular clusters are still
forming in the local universe*

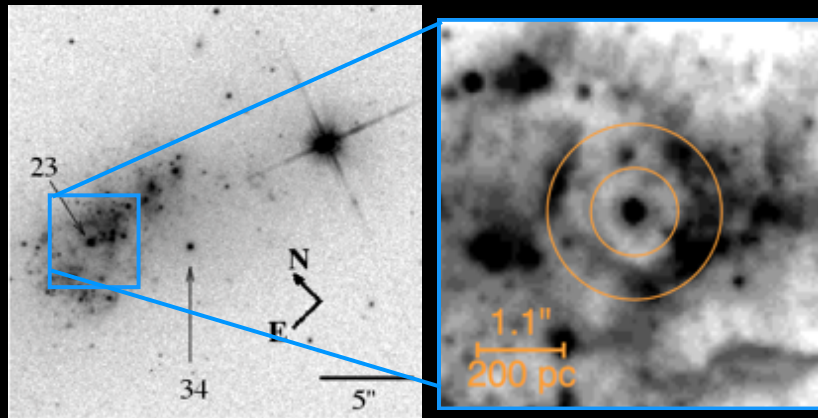
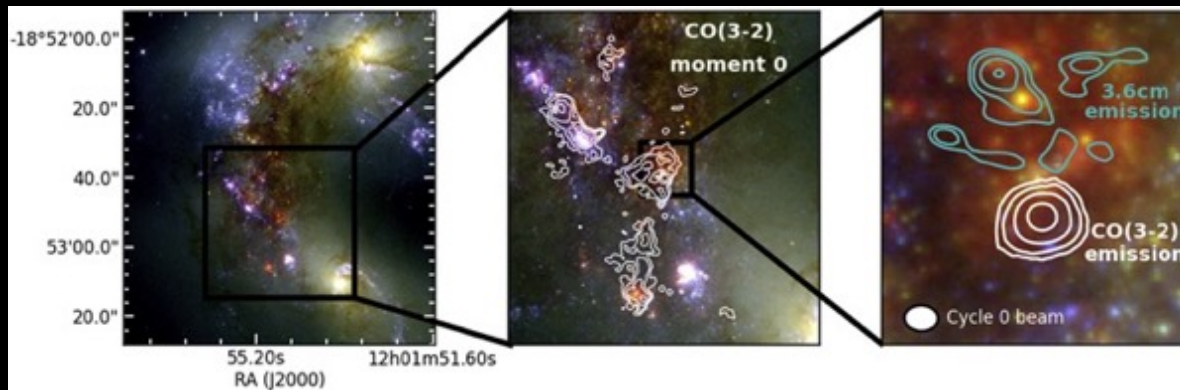


Antennae colliding
galaxies - HST

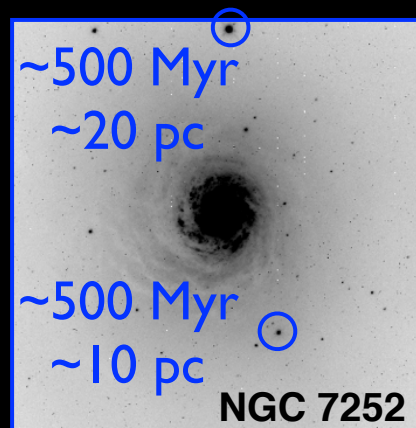
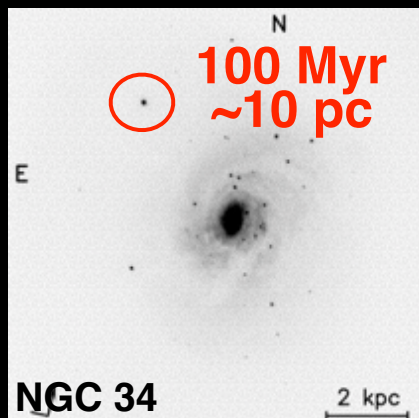
few million M_{sun}
 ~ 7 Myr



10 million solar mass clusters



We have examples
of $10^7 M_{\odot}$ at all
stages of their
evolution



What do we know about YMCs?

- ❖ Power-law mass function (index of -2 from ~ 100 to 10^8 Msun)
- ❖ High pressure environments conducive to forming high mass clusters
Elmegreen & Efremov 1998, Kruijssen 2015
- ❖ Single bursts of star-formation (no significant age spreads)
Bastian et al. 2013, Cabrera-Ziri et al. 2014, 2016
- ❖ Extremely efficient at removing gas/dust within them (< 4 Myr)
Hollyhead et al. 2015, Longmore 2015, Cabrera-Ziri et al. 2015
- ❖ In tension with all predictions from multiple populations scenarios
Bastian & Lardo 2018 ARA&A

Models for multiple populations do not agree with observations of YMCs

Are multiple populations detected in YMCs?

- ✧ Clusters younger than ~ 2 Gyr, none spectroscopically detected

Mucciarelli et al. 2008, 2012, 2014

- ✧ YMCs show complex CMDs. However, not likely connected to multiple populations (instead stellar rotation is the cause)

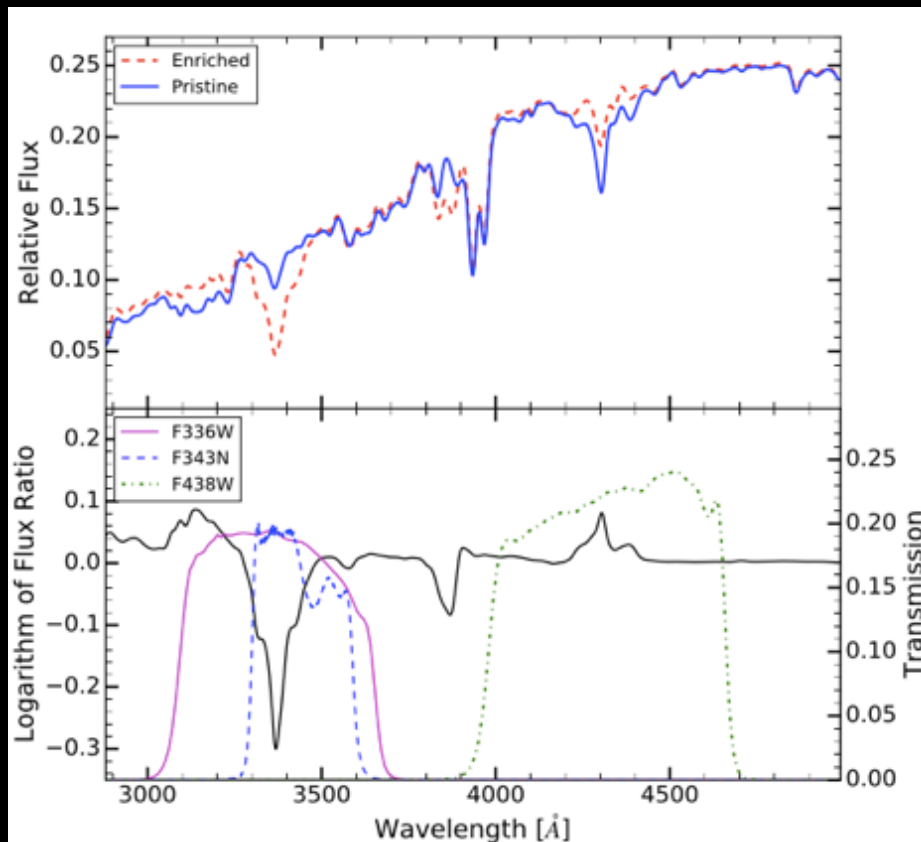
Milone et al. 2015, 2017; D'Antona et al. 2015, 2017; Bastian et al. 2015, 2017

- ✧ See upcoming Annual Review of Astronomy & Astrophysics

Bastian & Lardo 2018

Are Multiple Populations Restricted to Old GCs?

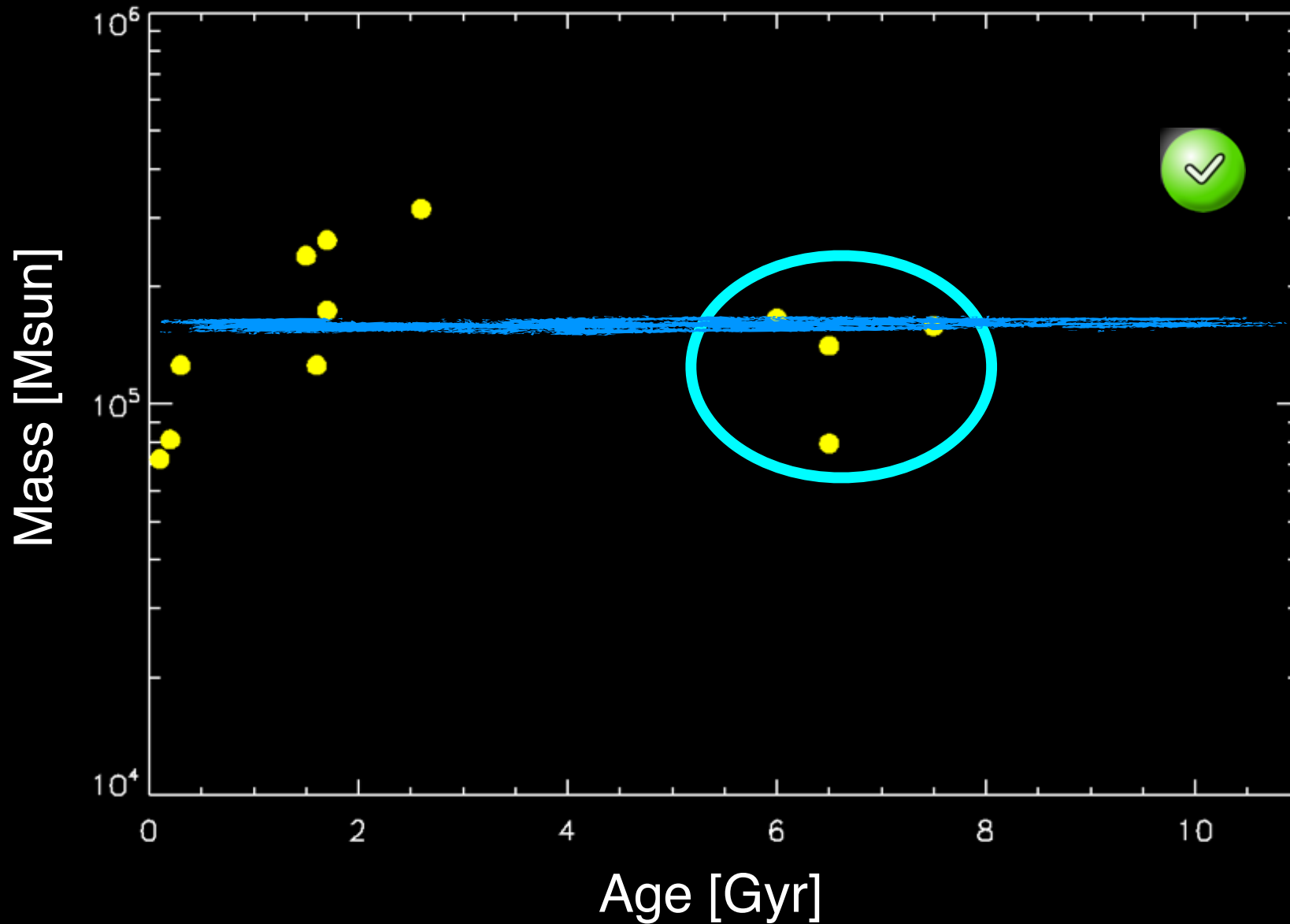
- ❖ HST study to search for multiple populations in massive ($>10^5 M_\odot$) clusters in the LMC/SMC, from 100 Myr - 11 Gyr
- ❖ VLT survey to search for N, C and O variations directly



The chosen filters pick up variations in N and O

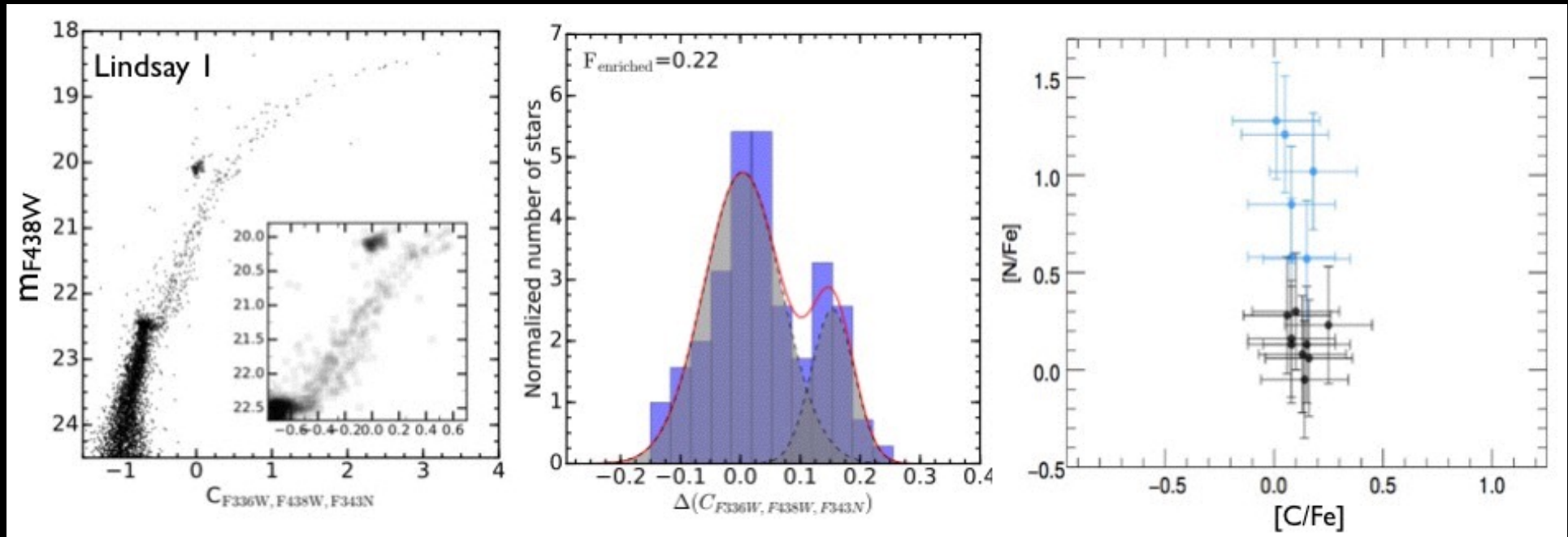
Niederhofer, NB, et al. 2016a,b
Hollyhead et al. 2016, 2017
Martocchia, NB et al. 2017

Are Multiple Populations Restricted to Old GCs?



Are Multiple Populations Restricted to Old GCs?

i.e. are YMCs and GCs really the same thing?



Lindsay I - SMC cluster

~ 7.5 Gyr

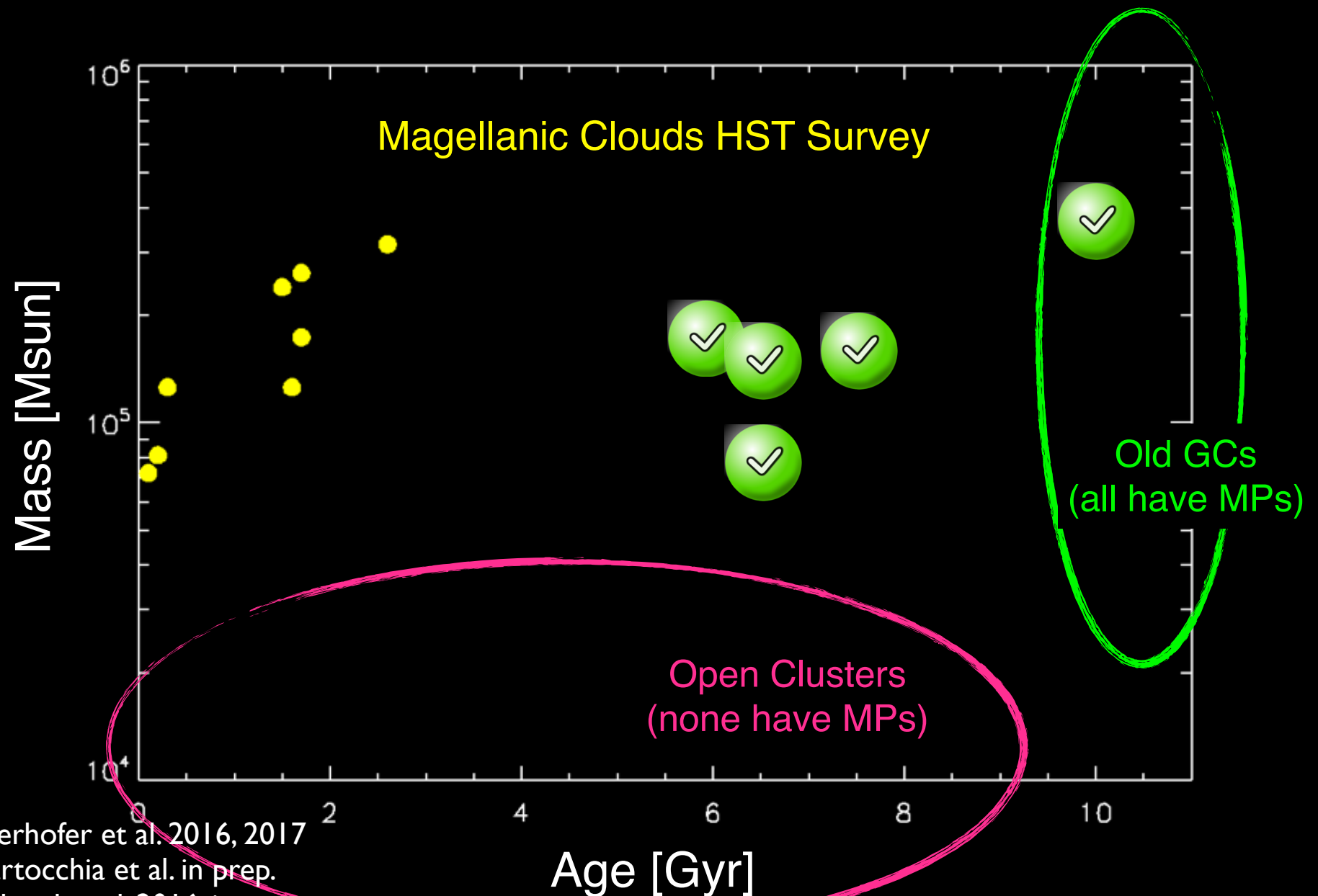
$\sim 10^5 M_{\odot}$

$Z_{\text{form}} \sim 1$

Hollyhead et al. 2016
Niederhofer, NB et al. 2017



Are Multiple Populations Restricted to Old GCs?



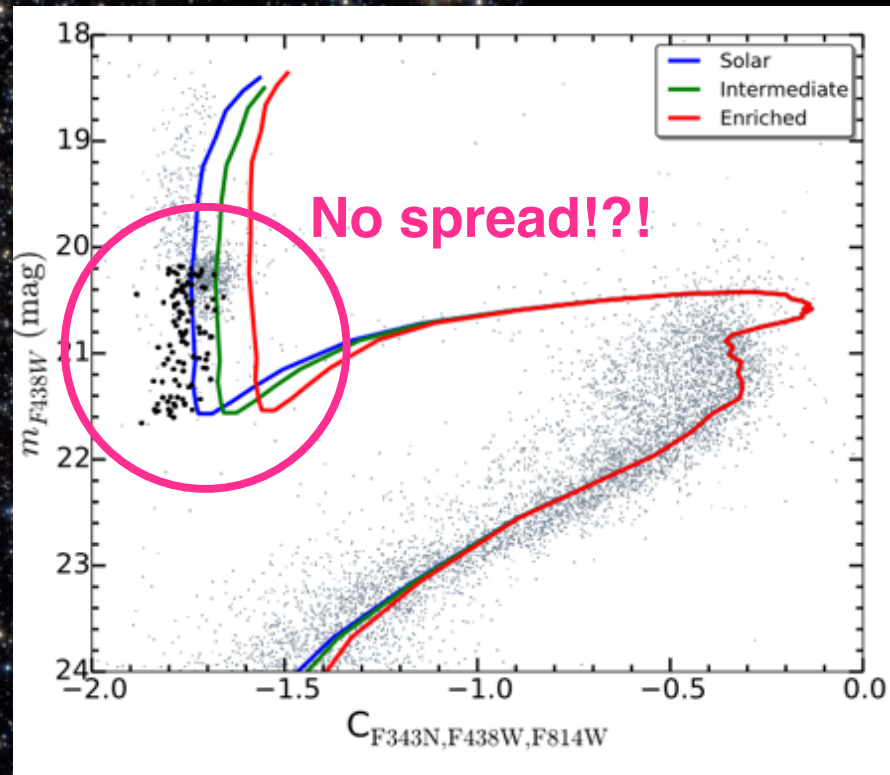
Niederhofer et al. 2016, 2017
Martocchia et al. in prep.
Hollyhead et al. 2016, in prep.

Are Multiple Populations Restricted to Old GCs?

NGC 419 - SMC cluster

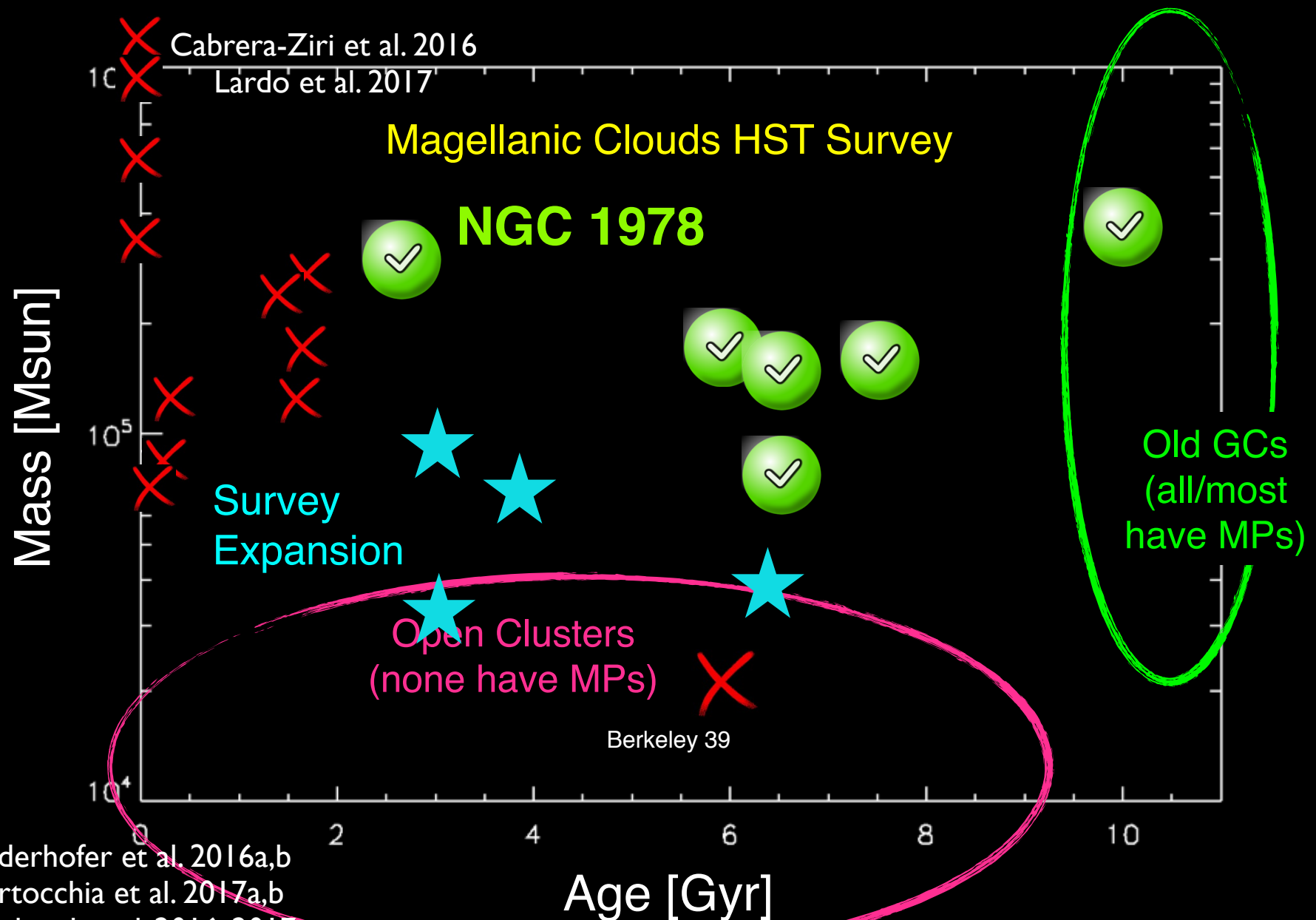
~1.5 Gyr

$\sim 2 \times 10^5 M_{\odot}$



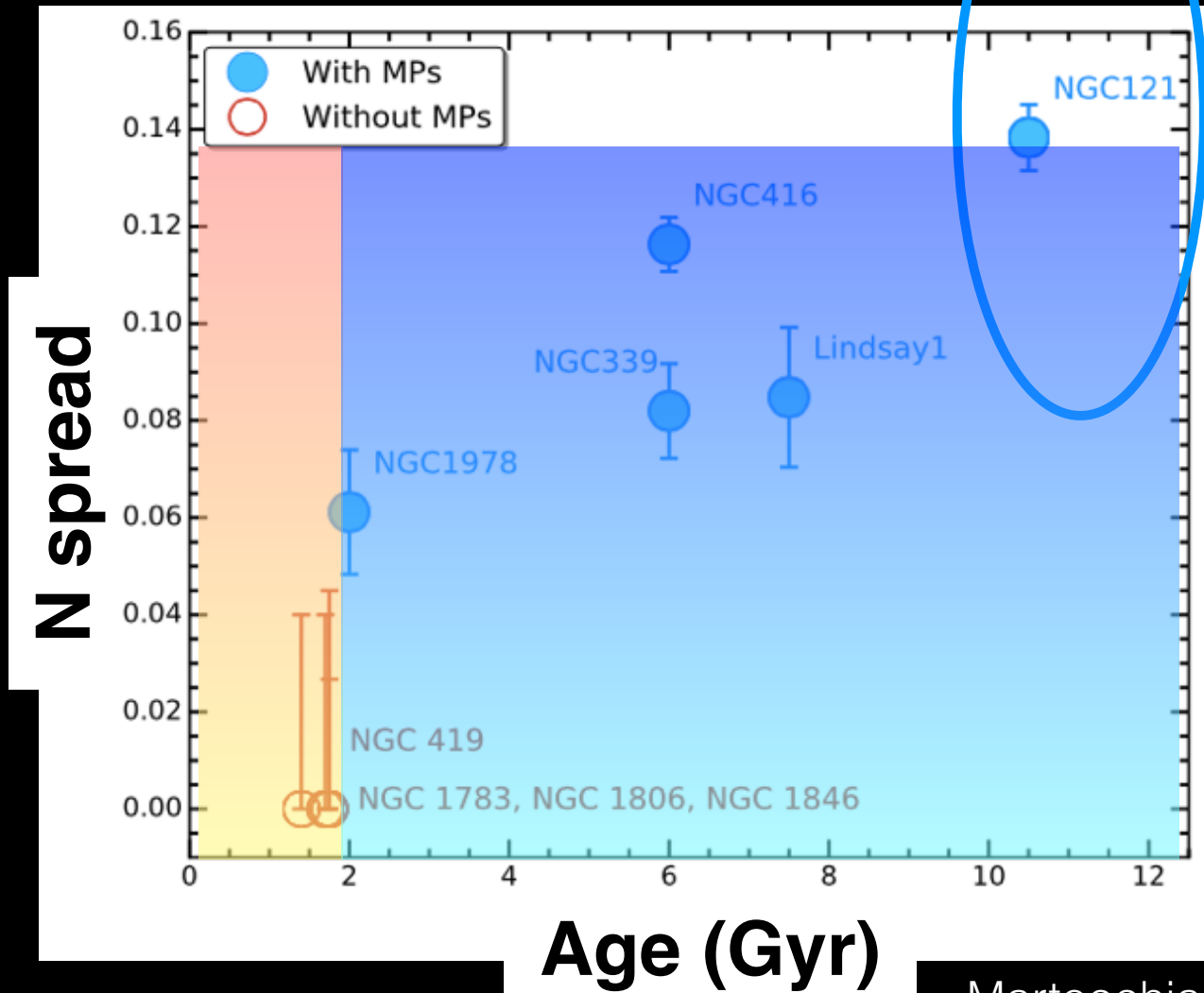
Martocchia, NB et al. 2017

Are Multiple Populations Restricted to Old GCs?



Role of cluster age

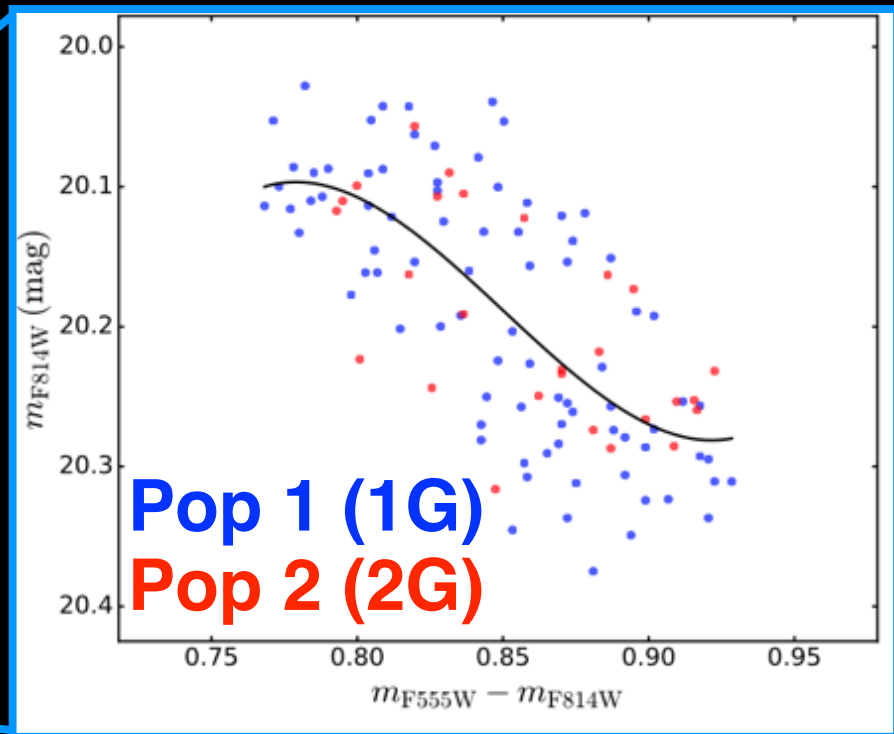
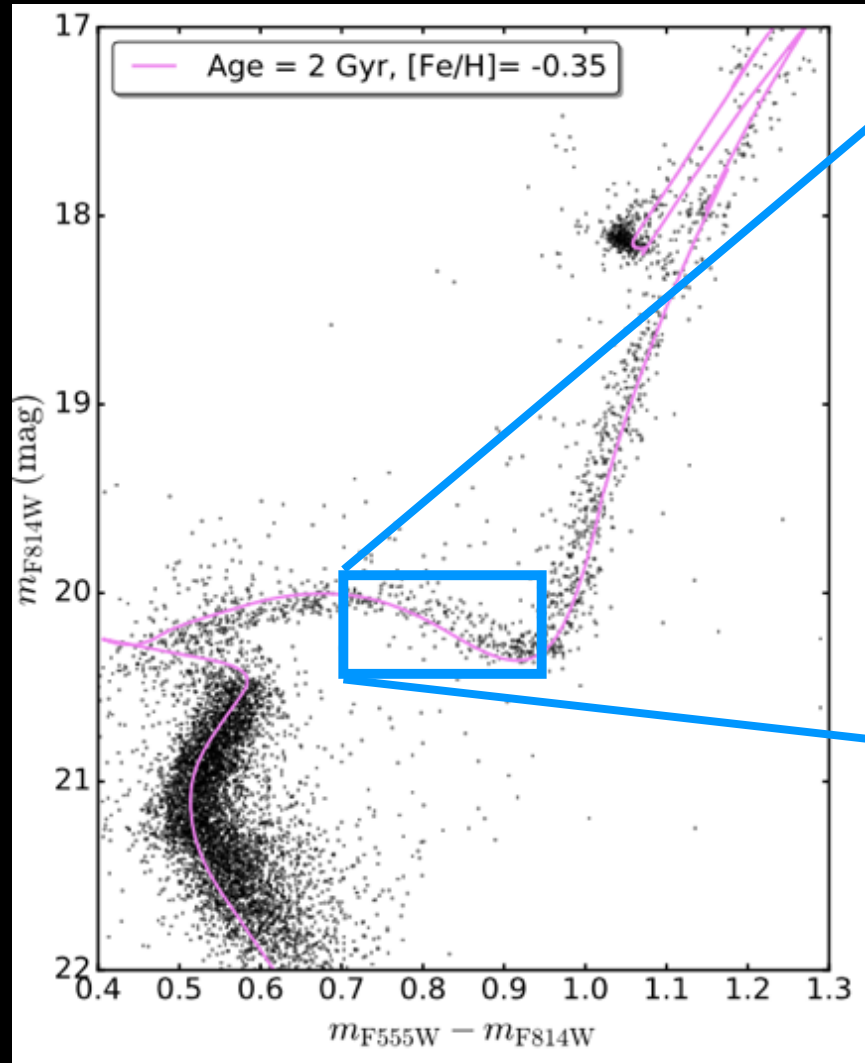
Milky Way
GCs



Martocchia et al. 2017a,b
Niederhofer et al. 2017a,b
Hollyhead et al. 2017

NGC 1978

Direct Constraints on Age

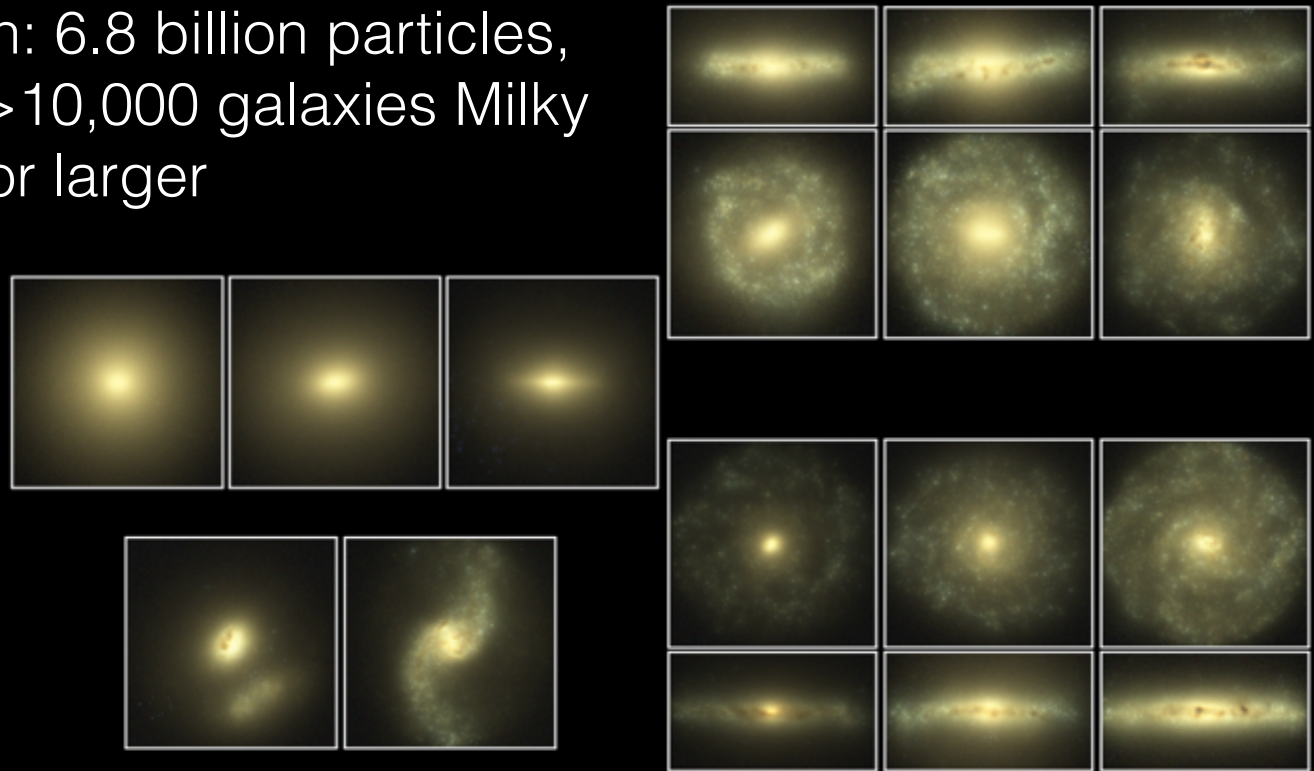


$$\Delta(\text{age}) = 6 \pm 17 \text{ Myr}$$

	YMCs (\$5.6)	Abundances (\$5.1)	Variety/stochasticity (\$5.1)	Mass budget (\$5.1.3)	He spread correlation with GC mass (\$5.4)	F ^{enriched} with GC mass (\$5.5)	Lack of trends with metallicity (\$5.5)	Age Dependence (\$2)	Li correlations (\$5.6.3)	(near) Ubiquity (\$5.1.2)	Discreteness (\$2)	Mg-Al (\$5.1.5)
AGB	✗	✗*	✗	✗	✗*	✗*	✗	✗	✗*	✗	✗*	✓*
FRMS	✗	✗*	✗	✗	✗	✗	✓	✗	✗	✗	✗	✗*
VMS	✓	✗*	?	✗	✓*	✓*	?	✗	✗	✗	✗	✗*
EDA	✓*	✗*	✗	✗	✗	✗	✓	✗	✗	✗	✗	✗
Reverse Order	✗	✗*	?	✓*	✗	✗	?	✗	✗	✗	✗	✗
eSF Period	✓*	✗*	✗	✗	✗	✗	?	✗	✗	✗	✗	✗

E-MOSAICS: MOdelling Star cluster system Assembly In Cosmological Simulations with the EAGLE simulations

EAGLE: Largest run: 6.8 billion particles,
100 Mpc box, with $>10,000$ galaxies Milky
Way or larger



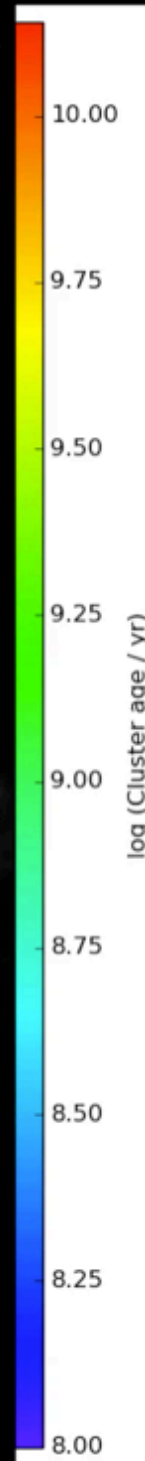
E-MOSAICS: MOdelling Star cluster system Assembly In Cosmological Simulations with the EAGLE simulations

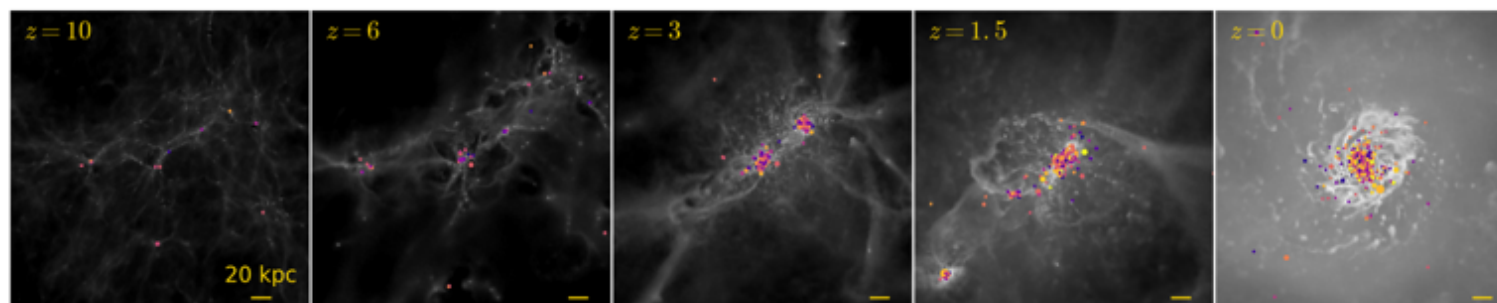
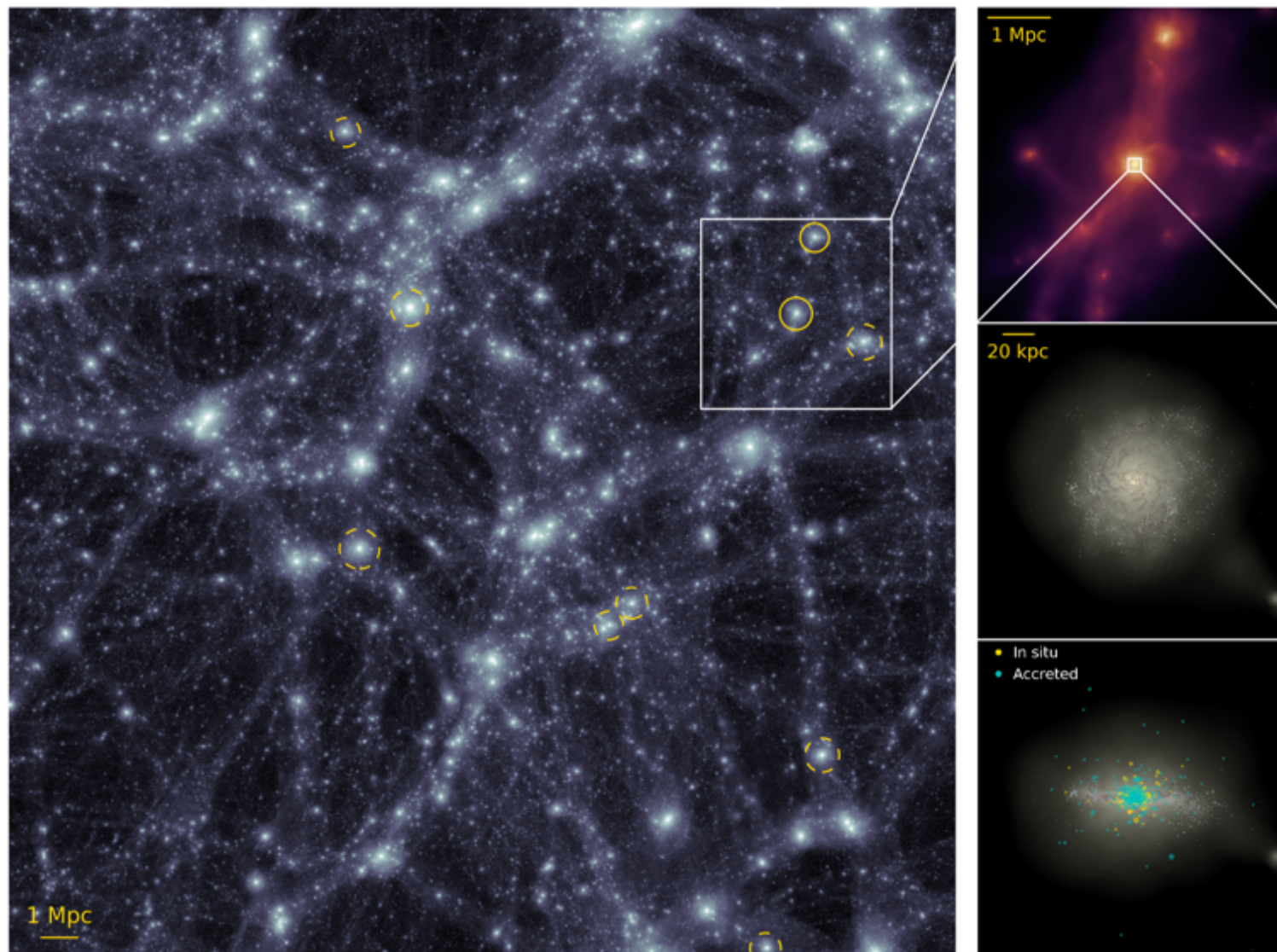
Simulation by Joel Pfeffer, Diederik Kruijssen, Rob Crain, Nate Bastian, Joop Schaye

Gas density and star cluster age

$z = 20.0$
 $t = 0.2 \text{ Gyr}$
 $L = 6.3 \text{ cMpc}$

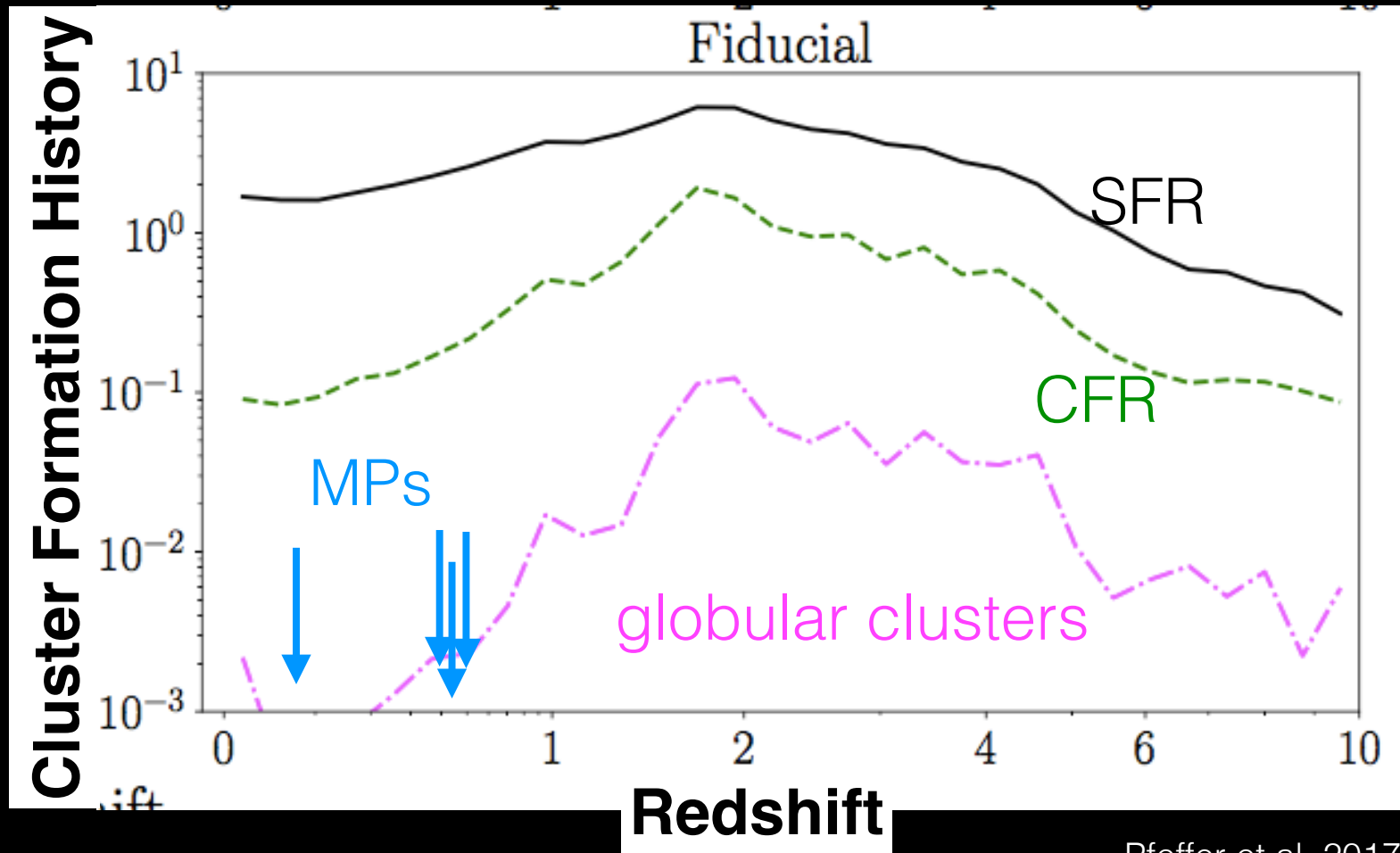
Pfeffer, Kruijssen, Crain & Bastian 2017





Pfeffer et al. 2017

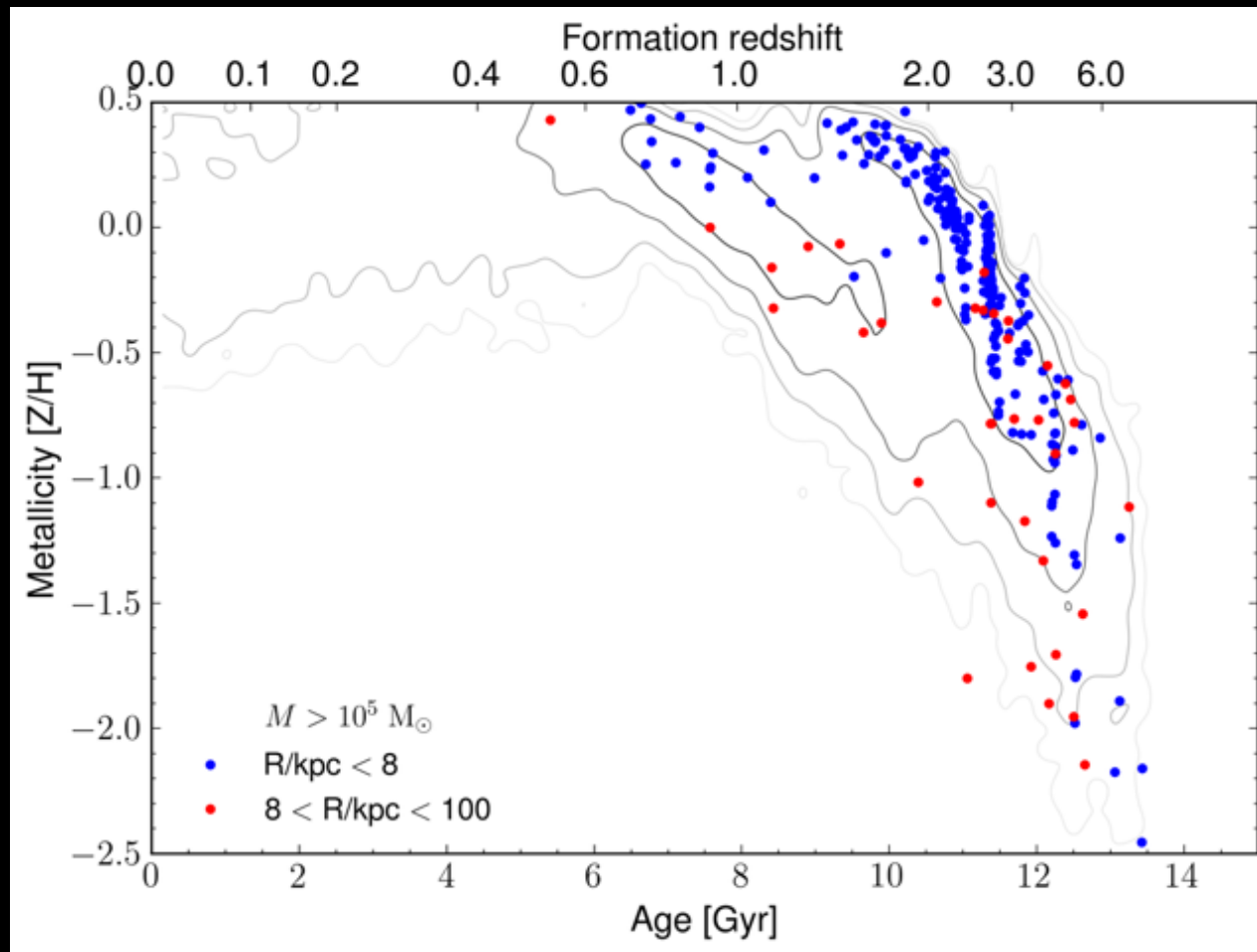
GC and YMC Formation in Cosmological Simulations



Pfeffer et al. 2017

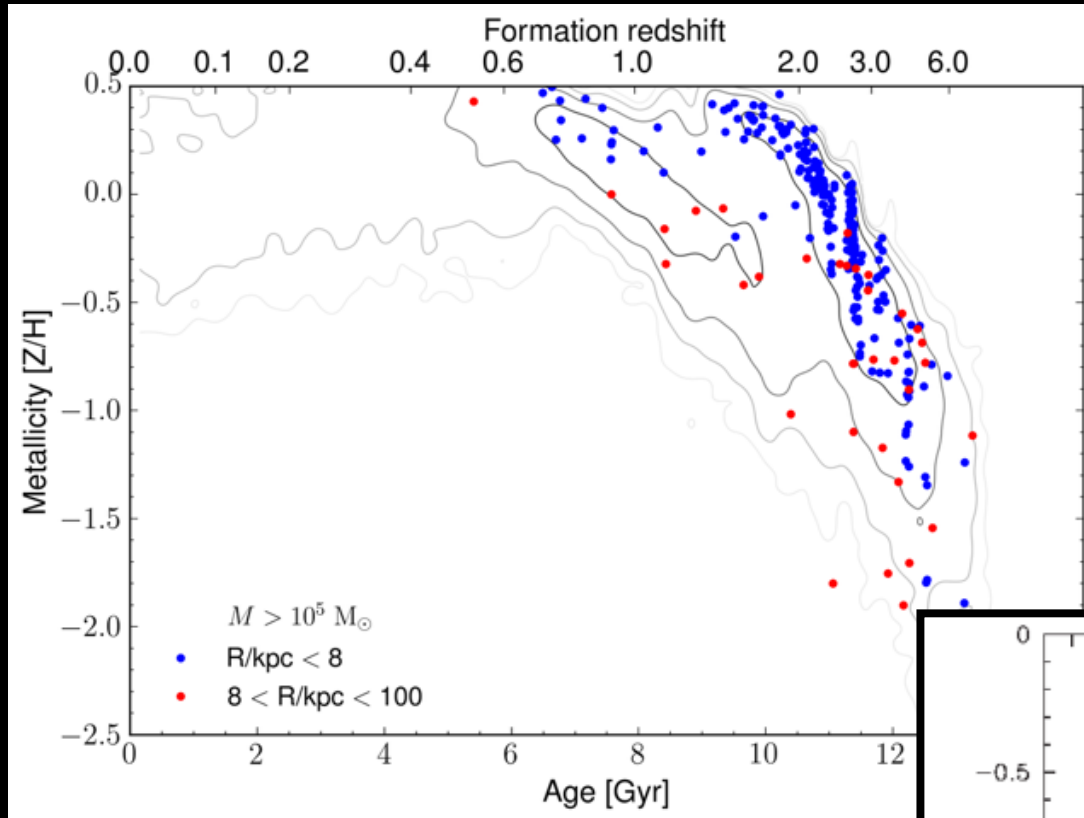
Reina-Campos et al. in prep.

GC and YMC Formation in Cosmological Simulations



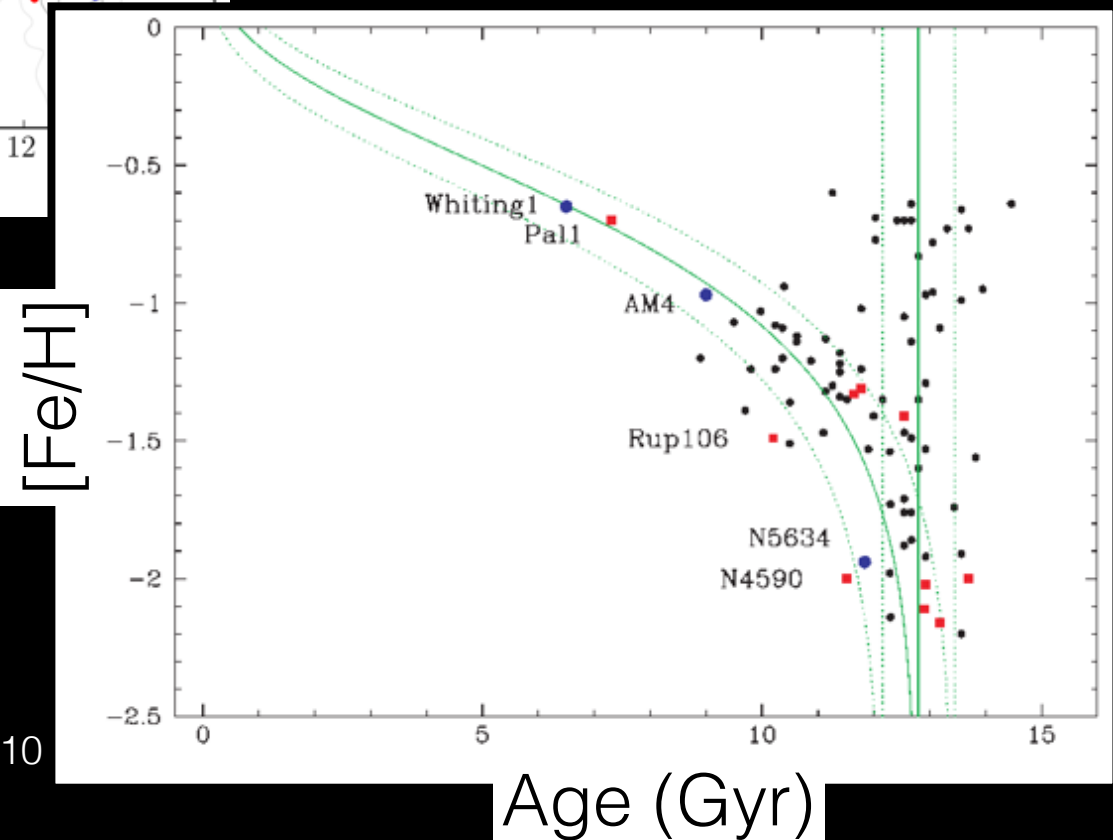
Kruijssen et al. in prep.

Age-metallicity distribution gives assembly history



Kruijssen et al. in prep.

Forbes & Bridges 2010



E-MOSAICS: MOdelling Star cluster system Assembly In Cosmological Simulations with the EAGLE simulations

Simulation by Joel Pfeffer, Diederik Kruijssen, Rob Crain, Nate Bastian, Joop Schaye

Gas density and star cluster age

$z = 20.0$
 $t = 0.2 \text{ Gyr}$
 $L = 6.3 \text{ cMpc}$



Pfeffer et al. 2017